Observations and Simulations of the Stable Water Isotope Signature of Shallow Trade Wind Clouds

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I. Introduction

Shallow trade-wind clouds are a source of uncertainty in climate projections. The coupling between clouds, convection and circulation is still largely a mystery. Here, we present airborne stable water vapour isotope measurements from EUREC4A(1) and complementary simulations using COSMOiso(2), which characterise the water vapour at cloud base and possibly point towards different processes depending on the prevailing cloud organisation pattern⁽³⁾.



comparable to the flight

segments cloud layer (lowest

water > 10⁻⁵ kg kg⁻¹), subcloud

surface and cloud lauer), and

above inversion (altitude taken from ATR flight track).

Shown are vertical profiles

of the 223 grid points inside the red box (Fig. 2.2a), the

median (horizontal lines;

identified layers for the 117

grid points with liquid clouds

10-90 percentile range (shading) of the three

(exemplary time step).

used for data extraction), and

grid cell with liquid cloud

layer (mid-level between

- O1 Does COSMO₁₅₀ reproduce the observed isotope signals and cloud organisation patterns?
- O2 Are the archetypal mesoscale cloud organisation patterns in the Trades associated with different isotope signals?
- O3 Is the transport history of importance for the air parcels' isotope signature when arriving in the subtropics?

2. Measurements onboard the ATR-42⁽¹⁾

- 19 flights (each ~4.5h) upstream of Barbados on II days between 25 Jan and 13 Feb 2020
- different flight segments with most time spent at cloud base in the so-called *cloud layer* (rectangular flight pattern)

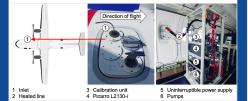


Figure 2.1 Measurement setup during EUREC4A with the laser spectrometer Picarro L2130-i deployed onboard the French aircraft ATR-42.

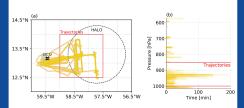
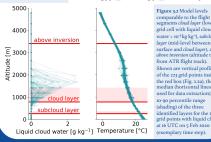


Figure 2.2 Flight path of 19 ATR flights overlayed (yellow). The area from which back-trajectories were calculated (red box/layer). The location of the Barbados Cloud Observatory (BCO) on Barbados and the circle flown by HALO(1) shown as black cross and circle, respectively

3. Simulations with COSMOiso

- North Atlantic domain (Fig. 3.1): initial and lateral boundary conditions every 6h from ECHAM6-wiso(4) (wind, PS, T nudged to ERA5); horizontal winds at p ≤ 850 hPa nudged to ECHAM6-wiso (wavenumbers ≤ 5)
- hourly output (explicit convection) on 40 vertical levels with 10 x 10 km horizontal resolution from 16 Dec 2019 - 20 Feb 2020
- inside red box (Fig. 2.2): data extraction at three distinct layers (Fig. 3.2) and LAGRANTO(5,6) back-trajectories with 20 km horizontal and 15 hPa (1000-850 hPa) vertical spacing





4. Model validation

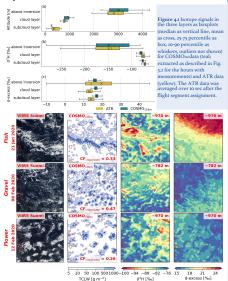
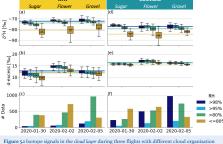
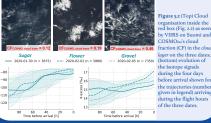


Figure 4.2 Cloud organisation (50-60°W, 10-20°N) shown by (left) Suomi, (centre left) COSMOise otal liquid cloud water and cloud fraction at cloud base (grid cells with liquid cloud water > 10.5 kg g-4), and (centre right/right) δ2H/d-excess in the cloud layer (determined as in Fig. 3.2 using all grid points in the domain) in COSMO₈₀. The dates are not identical to the ones shown in Box 5

5. Isotopes and cloud organisation



patterns (Sugar, Flower, Gravel)(3) for cloudy (RH>80%) and clear-sky (RH<=80%) environments in the (left) ATR and (right) COSMOso data. Boxplots as in Fig. 4.1; COSMOsodata extracted as described in Fig. 2.2 for the hours with measurements. (Bottom) data count in seconds for the ATR and in number of grid points during the ATR flight time for the COSMOso data. The ATR data wa iveraged over to sec after the assignment to the four RH categories



by VIIRS on Suomi and COSMOw's cloud fraction (CF) in the cloud lauer on the three dates bottom) evolution of the isotope signals during the four days before arrival shown for the trajectories (number during the flight hours of the three dates

6. Conclusion

- CI COSMO_{iso} underestimates the variability of the observed isotope signals, but qualitatively captures the gradients
- across different layers (Fig. 4.1). Observed mesoscale cloud organisations are fairly well reproduced in the simulations (Fig. 4.2). For δ^2 H, the in-flight variability is larger than the one between flights. For d-excess it is the other way around (Fig. 5.1). It is yet unclear how the isotopic signature relates to the mesoscale cloud organisation.
- While the δ^2 H signal forms shortly (~20 h) before arrival and is most likely controlled by the cloud environment, the d-excess anomaly appears earlier (~50 h before arrival) and is transported towards Barbados (Fig. 5.2).

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